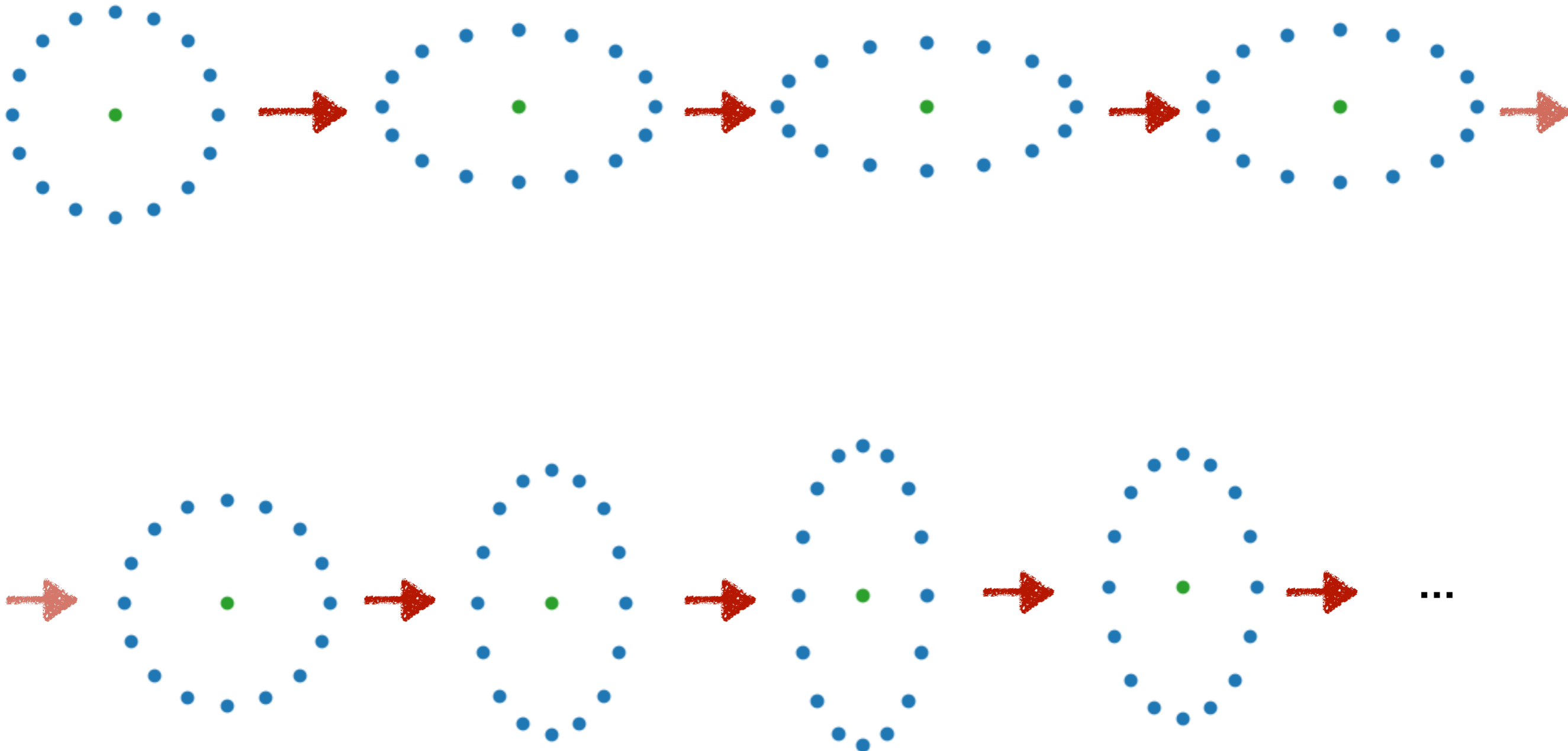


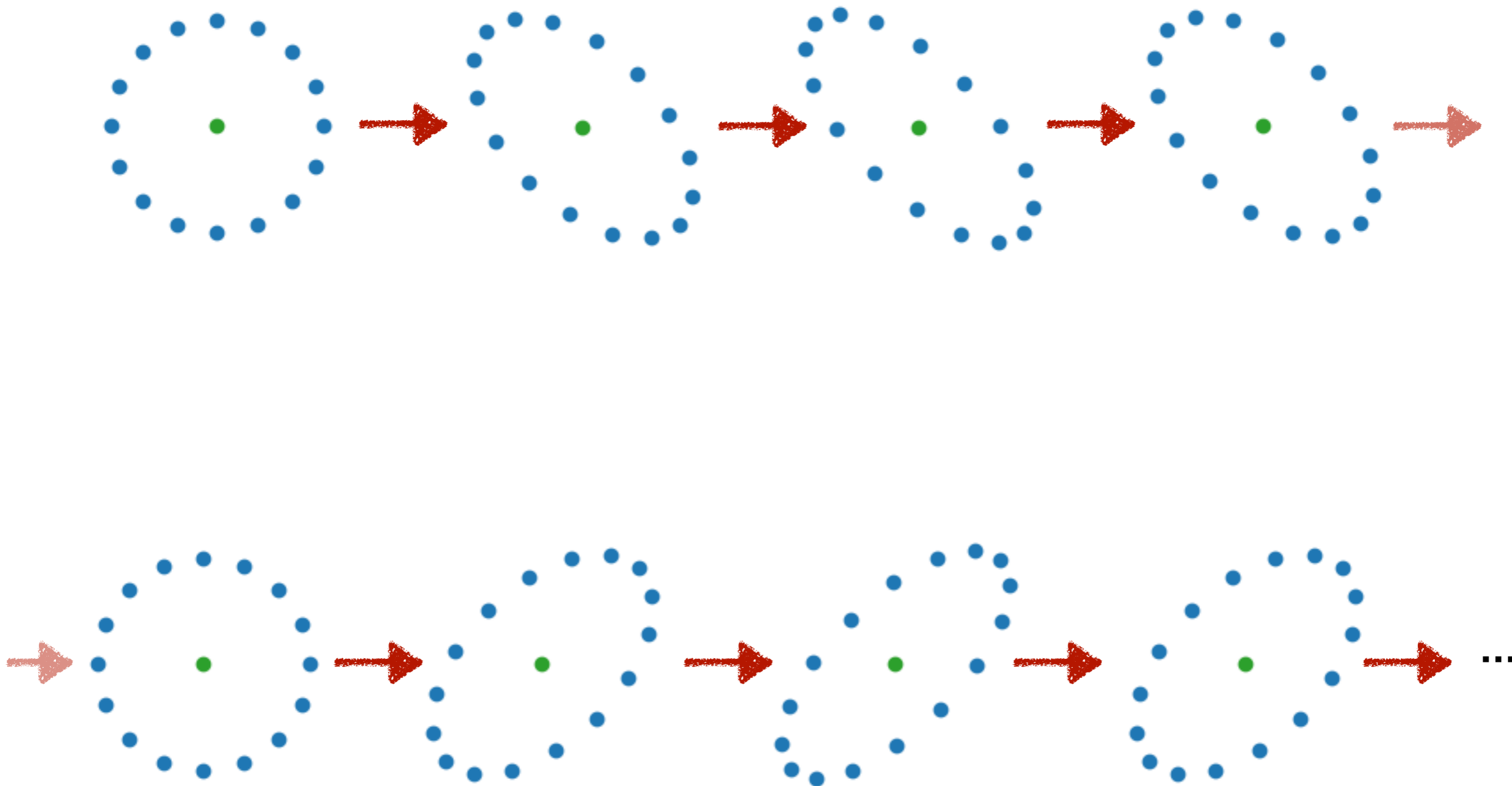
Gravitational waves

+ mode



Gravitational waves

X mode



Gravitational waves

Energy flux of a plane wave

Gravitational waves

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Fundamental problem with a local energy density of gravitational field

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...but for GW at perturbative level there's a reasonable expression

$$F = \frac{h_0^2 \omega^2}{32\pi G} = \frac{\pi f^2}{8G}$$

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Ch. 7.6, p. 307

effective stress-energy tensor for gravitational waves $t_{\mu\nu} = \frac{1}{32\pi G} \left\langle \partial_\mu h_{\alpha\beta}^{TT} \partial_\nu h_{TT}^{\alpha\beta} \right\rangle$

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1. Vacuum Einstein equations up to 2nd order

$$g_{\mu\nu} = \eta_{\mu\nu} + h_{\mu\nu}^{(1)} + h_{\mu\nu}^{(2)}$$

$$R_{\mu\nu}^{(1)} [h_{\mu\nu}^{(1)}] = 0$$

$$R_{\mu\nu}^{(1)} [h_{\mu\nu}^{(2)}] + R_{\mu\nu}^{(2)} [h_{\mu\nu}^{(1)}] = 0$$

$$R_{\mu\nu} = R_{\mu\nu}^{(1)} + R_{\mu\nu}^{(2)}$$

linearized
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quadratic term in
the expansion of Ricci

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2. Re-writing the equations for $h_{\mu\nu}^{(2)}$

$$R_{\mu\nu}^{(1)} [h_{\alpha\beta}^{(2)}] - \frac{1}{2} \eta_{\mu\nu} \eta^{\kappa\lambda} R_{\kappa\lambda}^{(1)} [h_{\alpha\beta}^{(2)}] = 8\pi G t_{\mu\nu}^{\text{loc}}$$

$$t_{\mu\nu} = -\frac{1}{8\pi G} \left(R_{\mu\nu}^{(2)} [h_{\alpha\beta}^{(1)}] - \frac{1}{2} \eta_{\mu\nu} \eta^{\kappa\lambda} R_{\kappa\lambda}^{(2)} [h_{\alpha\beta}^{(1)}] \right)$$

Gravity gravitates!

**Higher order terms in metric „feel”
the gravity of a GW**

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- local
- conserved: $\partial_{\mu} t_{\text{loc}}^{\mu\nu} = 0$
- gauge-dependent

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3. Average over many wavelengths

$$\langle \dots \rangle = V_D^{-1} \int_D (\dots) d^4x$$

$$t_{\mu\nu} = \langle t_{\mu\nu}^{\text{loc}} \rangle$$

$$t_{\mu\nu} = \frac{1}{32\pi G} \left\langle \partial_{\mu} h_{\alpha\beta}^{TT} \partial_{\nu} h_{TT}^{\alpha\beta} \right\rangle$$

$$t^{0z} = F = \frac{h_0^2 \omega^2}{32\pi G}$$